

Large lepton mixings from continuous symmetries

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ABSTRACT: Within the broad context of quark-lepton unification, we investigate the implications of broken continuous family symmetries which result from requiring that in the limit of exact symmetry, the Dirac mass matrices yield hierarchical masses for the quarks and charged leptons, but lead to *degenerate* light neutrino masses as a consequence of the seesaw mechanism, without requiring hierarchical right-handed neutrino mass terms. Quark mixing is then naturally small and proportional to the size of the perturbation, but lepton mixing is large as a result of degenerate perturbation theory, shifted from maximal mixing by the size of the perturbation. Within this approach, we study an illustrative two-family prototype model with an $SO(2)$ family symmetry, and discuss extensions to three-family models.

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1. Introduction

Neutrino oscillation experiments have confirmed that neutrinos are massive and lepton mixing is observable. This result, which is the most important recent experimental result in particle physics, has brought a new level of confusion and excitement to the already puzzling area of flavor physics. Although the triumph of the Standard Model (and its extension to grand unified theories) suggests a unified view of the quarks and leptons in terms of the gauge interactions, quarks and leptons display striking differences in terms of their masses and mixings.

In the quark sector, the Cabibbo-Kobayashi-Maskawa mixing matrix is approximately the identity matrix, and the masses are equally spaced on the logarithmic scale. Parametrized in terms of the Cabibbo angle $\lambda_c \equiv \sin \theta_c = 0.22$ ($\theta_c \simeq 13^\circ$), the quark mass ratios are

$$\frac{m_u}{m_t} \sim \lambda_c^8, \quad \frac{m_c}{m_t} \sim \lambda_c^4, \quad \frac{m_d}{m_b} \sim \lambda_c^4, \quad \frac{m_s}{m_b} \sim \lambda_c^2. \quad (1.1)$$

As a result, perturbing about the limit of vanishing Cabibbo angle is a popular approach to understanding the flavor problem.

In the lepton sector, the situation is quite different. The Maki-Nakagawa-Sakata-Pontecorvo (MNSP) [1] lepton mixing matrix is experimentally known [2–4] to be of the form (neglecting phases):

$$\mathcal{U}_{\text{MNSP}} \approx \mathcal{R}_1(\theta_\oplus) \mathcal{R}_3(\theta_\odot), \quad (1.2)$$

in which $\mathcal{R}_a(\theta)$ denotes a rotation matrix about the a -axis. The atmospheric and solar mixing angles are large:

$$\theta_\oplus \approx 45^\circ \pm 10^\circ, \quad \theta_\odot \approx 34^\circ \pm 2^\circ. \quad (1.3)$$

Equally puzzling is the fact that oscillations are predominantly among two neutrinos; the third mixing angle θ_{13} is bounded to be at most of Cabibbo size [4]. Turning now to lepton masses, the charged lepton masses display hierarchies

$$\frac{m_e}{m_\tau} \sim \lambda_c^5, \quad \frac{m_\mu}{m_\tau} \sim \lambda_c^2, \quad (1.4)$$

but the same is not obvious for the neutrinos. Oscillation data has indicated that the ratio of the observed solar and atmospheric mass-squared differences can be expressed in terms of λ_c as follows:

$$\frac{\Delta m_{\odot}^2}{\Delta m_{\oplus}^2} \simeq \frac{8 \times 10^{-5} \text{ eV}^2}{2 \times 10^{-3} \text{ eV}^2} \sim \lambda_c^2. \quad (1.5)$$

Assuming only three active neutrinos oscillate, with masses $m_{1,2,3}$, there are several possible neutrino mass patterns, all of which display milder hierarchies than those of the charged fermions:

- Normal hierarchy ($m_1 < m_2 \ll m_3$):

$$\frac{m_1}{m_3} \sim \sqrt{\left(\frac{m_2}{m_3}\right)^2 - \lambda_c^2}, \quad \frac{m_2}{m_3} \sim \lambda_c, \quad (1.6)$$

- Inverted hierarchy ($m_3 \ll m_1 \sim m_2$):

$$\frac{m_1}{m_2} \sim 1 + \lambda_c^2, \quad \frac{m_2}{m_3} \lesssim \lambda_c. \quad (1.7)$$

- Quasi-degenerate ($m_1 \sim m_2 \sim m_3$):

$$\frac{m_1}{m_3} \sim \frac{m_2}{m_3} \sim 1 + \lambda_c, \quad \frac{m_1}{m_2} \sim 1 + \lambda_c^3. \quad (1.8)$$

The data have challenged us to understand the origin of the observed discrepancies between the quark and lepton sectors. This question is particularly intriguing within the context of quark-lepton unification [5], for which all available data can be synthesized in the search for a compelling flavor theory. The lepton data has led to a recent renaissance in flavor model building, with many examples in the literature (see the reviews [6] and references therein).

The first step in addressing the flavor puzzle is to seek a reasonable theoretical starting point for building flavor theories. To this end, quark-lepton unification compels us to consider the effects of small (e.g. Cabibbo sized) perturbations in the lepton sector. For the mixings, θ_{13} is akin to quark angles (deviation from zero), while for the solar and atmospheric mixings, small parameters would appear as *deviations* from large initial values. The measured angles are thus “hazed” away from starting values which assume values indicative of flavor symmetries [7–9] (in other approaches θ_{\oplus} and θ_{\odot} take their *central* experimental values as indicative of discrete flavor symmetries [10]).¹ Small parameters may govern a subset of the neutrino masses or their deviations from a common mass. However, given the large lepton mixings, a natural theoretical starting point is neutrino mass degeneracy, since large angles can arise from small perturbations about degenerate structures.

¹Alternatively, one can give up the study of flavor physics, appeal to some landscape, or like number theorists of old, ascribe magical meanings to the central values of the angles (our favorite is the Phidian angle (golden mean) $\pi/5 \approx 36^\circ$).

In that spirit, we investigate the implications of the family symmetries which emerge by *requiring* a theoretical starting point where in the limit of exact symmetry, the Dirac ($\Delta I_w = 1/2$) and Majorana ($\Delta I_w = 0$) mass matrices yield hierarchical masses for the charged fermions and degenerate masses for *both* the light neutrinos obtained from the seesaw mechanism [11] and the heavy right-handed neutrinos. These requirements lead to specific constraints on the $\Delta I_w = 1/2$ sector which suggest features of the family symmetry. Note that the path to neutrino mass degeneracy differs here from that most often found in the literature (e.g. in single right-handed neutrino dominance models [12]), in which hierarchical right-handed neutrino masses offset the hierarchical Dirac masses. In its simplest implementation, an $SO(2)$ family symmetry results from these requirements. Large lepton mixing angles then arise from small perturbations in the neutrino sector, as expected from degenerate perturbation theory.

In this paper, after arguing that quasi-degenerate neutrinos are a reasonable starting point on both phenomenological and theoretical grounds, we outline our approach for an arbitrary number of families (assuming one right-handed neutrino per family). We then focus in detail on a prototype two-family model with the $SO(2)$ family symmetry which emerges naturally in our approach. In this scenario, the charge eigenstates are given by

$$\frac{1}{\sqrt{2}}(\psi \pm i\psi'),$$

where ψ and ψ' are the current eigenstates of the two families. The degeneracy of the light seesawed neutrinos is ultimately due to nontrivial phases in the couplings enforced by the symmetry. With symmetry breaking in either the $\Delta I_w = 0$ or the $\Delta I_w = 1/2$ neutrino sector, the degeneracy is lifted and the mixing angle is hazed away from 45° by the perturbation. We next discuss ways to extend the general approach to three-family models, and conclude by discussing avenues of future exploration.

2. Theoretical framework

The case for neutrino mass degeneracy

In our theoretical approach to the flavor puzzle, the discrepancies between the quark and lepton mixings are ultimately due to the theoretical requirement of neutrino mass degeneracy in the limit of an exact family symmetry. Hence, we first wish to emphasize that quasi-degenerate neutrinos are intriguing both from a phenomenological and a theoretical point of view. In particular, while the overall neutrino mass scale has been constrained from cosmology [13] (and, to a lesser extent, $0\nu\beta\beta$ [14]), quasi-degenerate scenarios are consistent with the data (though with a reduced parameter space). The cosmological constraints yields the following bound on the sum of the neutrino masses [13]:

$$\sum_i m_i \lesssim 0.2 - 2 \text{ eV}. \tag{2.1}$$

The range reflects the dependence of the bound on priors and the data sets used for the fits. Assuming a conservative estimate of the bound of 0.9 eV, the overall neutrino mass

scale m_0 for quasi-degenerate models is $m_0 \lesssim 0.3 \text{ eV}$. Such scenarios have the advantage that they will be probed by neutrinoless double beta decay experiments [14] and direct beta decay experiments [15].

From a theoretical standpoint, one may worry that if the experimental bound m_0 is not much larger than the observed mass splittings, the ratio of the mass-splittings to the overall scale is not a reliable expansion parameter. This observation, which would naively appear to disfavor degenerate neutrinos as a theoretical starting point, can be countered in two ways. First, note that current bounds do not preclude scenarios in which this ratio is a small parameter. To see this more clearly, let us parametrize the masses as follows:

$$m_1 = m_0 - |\Delta_\ominus|, \quad m_2 = m_0, \quad m_3 = m_0 + \Delta_\oplus, \quad (2.2)$$

in which Δ_\oplus can take either sign. The experimentally measured mass-squared differences are then given by

$$\begin{aligned} \Delta m_\oplus^2 &\equiv |m_3^2 - m_2^2| = (2m_0 + \Delta_\oplus)|\Delta_\oplus| \\ \Delta m_\ominus^2 &\equiv |m_2^2 - m_1^2| = (2m_0 - |\Delta_\ominus|)|\Delta_\ominus|. \end{aligned} \quad (2.3)$$

The ratios of the mass-splittings to the overall mass scale take the form

$$\begin{aligned} \frac{\Delta_\oplus}{m_0} &= \pm \sqrt{1 \pm \frac{\Delta m_\oplus^2}{m_0^2}} - 1 \\ \frac{|\Delta_\ominus|}{m_0} &= \pm \sqrt{1 - \frac{\Delta m_\ominus^2}{m_0^2}} + 1. \end{aligned} \quad (2.4)$$

These ratios can be large numbers if $m_0 \ll \sqrt{\Delta m_{(\text{exp})}^2}$; however, this only occurs when m_0 is more than an order of magnitude below the cosmological limit. (The solutions with $\mathcal{O}(1)$ ratios independently of m_0 correspond to negative neutrino masses when $\Delta m_{(\text{exp})}^2 \rightarrow 0$, and hence effectively cover the same regime.)

In addition, note that $\mathcal{O}(1)$ or larger ratios are not necessarily problematic for seesaw models based on small perturbations about degenerate structures. Since the seesaw is a convolution of the $\Delta I_w = 1/2$ and $\Delta I_w = 0$ mass matrices, we will see that small perturbations in the Majorana masses can be enhanced by the hierarchies in the Dirac mass terms without invalidating perturbation theory. In the two-family $SO(2)$ model discussed later in the paper, the Dirac masses are $m_{1,2} \sim e^{\pm\eta}$ and there is a small parameter $\delta \ll 1$ in the electroweak singlet sector, yet the mass-squared difference is $\sim \delta \cosh 2\eta$ and the shift in the mixing angle is $\sim \delta \sinh 2\eta$. Therefore, Δ_\oplus/m_0 and Δ_\ominus/m_0 need not be small even when there are small parameters in the seesaw matrix. This is a point that, to our knowledge, has not been fully appreciated in the literature.

Systematics of the approach

With these ideas in mind, we now outline our theoretical approach. The main idea is to study the symmetries which emerge upon imposing the theoretical requirement that

at tree level² the charged fermion masses are hierarchical, but the light neutrino masses emerge naturally through the seesaw mechanism, without strong hierarchies in the right-handed neutrino sector.³ These requirements lead to basis-dependent conditions on both the $\Delta I_w = 0$ and the $\Delta I_w = 1/2$ mass matrices, which can be expressed in the basis in which the charged currents are diagonal (the current basis) as follows:

- **$\Delta I_w = 0$ sector.** The Majorana mass matrix for the right-handed neutrinos is proportional to the identity matrix:

$$\mathcal{M}_{\text{Maj}} \propto 1. \tag{2.5}$$

Assuming one right-handed neutrino per family, the symmetry is $O(n)$ (for n families). If the number of right-handed neutrinos per family is doubled, as in E_6 models [19], the symmetry can be larger, e.g. $SU(n)$.

- **$\Delta I_w = 1/2$ sector.** Hierarchical Dirac mass eigenvalues require that the Dirac mass matrices \mathcal{M}_D obey the following condition:

$$\mathcal{M}_D \mathcal{M}_D^\dagger \neq 1. \tag{2.6}$$

We require that this condition holds for the Dirac mass matrices for both the charged and neutral fermions. Given the form of \mathcal{M}_{Maj} above and recalling the form of the seesaw matrix,

$$\mathcal{M}_\nu = \mathcal{M}_D \mathcal{M}_{\text{Maj}}^{-1} \mathcal{M}_D^T, \tag{2.7}$$

degenerate light neutrino masses require an additional condition⁴ on the neutral Dirac mass matrix (which may also hold for the charged fermions):

$$(\mathcal{M}_D \mathcal{M}_D^T)(\mathcal{M}_D \mathcal{M}_D^T)^\dagger \propto 1. \tag{2.8}$$

Restricting ourselves for now to scenarios with an equal number of right-handed and left-handed neutrinos, this condition can be reexpressed as in terms of the Hermitian combination $\mathcal{H} \equiv \mathcal{M}_D^\dagger \mathcal{M}_D$ as follows:

$$\mathcal{H} \mathcal{H}^T \propto 1. \tag{2.9}$$

Writing $\mathcal{H} = \mathcal{S} + i\mathcal{A}$, in which \mathcal{S} and \mathcal{A} are real symmetric and antisymmetric matrices, respectively, eq. (2.9) translates to the conditions

$$\mathcal{S}^2 + \mathcal{A}^2 \propto 1; \quad [\mathcal{A}, \mathcal{S}] = 0,$$

which also imply

$$[\mathcal{A}, \mathcal{H}] = 0.$$

\mathcal{A} is a real antisymmetric matrix and thus is in the Lie algebra of $O(n)$. Hence, \mathcal{H} commutes with at least one generator of $O(n)$, requiring an $SO(2)$ symmetry at least as part of the

²We focus here on tree-level predictions. In complete scenarios one would need to consider higher-dimensional operators a la Froggatt-Nielsen [16]; we relegate this for future study.

³It is worth noting, however, that some hierarchy in the right-handed neutrino sector may be advantageous for leptogenesis [17] (but see, however, [18]).

⁴This condition holds because the observable quantities are the absolute values of the mass eigenvalues, and encompasses the special case in which $\mathcal{M}_D \mathcal{M}_D^T \propto 1$.

family symmetry of the theory. Eq. (2.9) also has implications for the Dirac mass hierarchy; the eigenvalues of H are either (up to an overall scale) 1 or mutually reciprocal pairs $e^{\pm\eta}$. For odd n , the middle Dirac mass eigenvalue is the geometric mean of the highest and lowest eigenvalues.⁵

The above discussion did not specify the number of families. Beginning with the case of two families for simplicity, we note that the unique solution for \mathcal{H} is

$$\mathcal{A} = c_A i\sigma_2, \quad \mathcal{S} = c_S 1, \tag{2.10}$$

where $c_{A,S}$ are arbitrary real constants. Writing \mathcal{M}_D as a sum of group elements T_i with complex coefficients ω_i , such that

$$\mathcal{M}_D = \omega_1 T_1 + \omega_2 T_2, \tag{2.11}$$

one solution is $T_1 = 1$ and $T_2 = i\sigma_2$, with $c_S = \omega_1^2 + \omega_2^2$, and $c_A = -2\text{Im}(\omega_1\omega_2^*)$.⁶ At least one complex ω_i is required, or the hierarchical structure of \mathcal{M}_D is lost (the eigenvalues form a mutually reciprocal pair). We will present explicit two-family $SO(2)$ models with these properties in section 3.

For three families, a generalization to $O(3)$ is suggested by the requirement on the $\Delta I_w = 0$ mass terms as given in eq. (2.5). However, the requirement that \mathcal{H} commutes with all of the generators of $O(3)$ implies that \mathcal{H} must be proportional to the identity, and hence there is no hierarchy in the Dirac mass eigenvalues. Therefore, an $O(3)$ symmetry in the $\Delta I_w = 1/2$ sector does not satisfy the constraints of our approach. For another way to see this, let us write \mathcal{M}_D once again as a sum of group elements,

$$\mathcal{M}_D = \omega_1 T_1 + \omega_2 T_2 + \omega_3 T_3, \tag{2.12}$$

and require that $\mathcal{M}_D \mathcal{M}_D^T \propto 1$. For $O(n)$, $T_i T_i^T = 1$, leading to the conditions

$$\omega_1 \omega_2 (T_1 T_2^T + T_2 T_1^T) + \omega_1 \omega_3 (T_1 T_3^T + T_3 T_1^T) + \omega_2 \omega_3 (T_2 T_3^T + T_3 T_2^T) = 0. \tag{2.13}$$

For $O(3)$, these conditions cannot be satisfied by inspection. One T_k must be chosen such that $T_i T_k^T = 0$ for $i \neq k$, corresponding to an $O(2)$ subgroup.

For $n > 3$, in principle it is possible to find at least one other generator of $O(n)$ which commutes with \mathcal{A} and some \mathcal{S} 's. We will continue the discussion of how to generalize the approach to three and more families in section 4.

3. Two family models

We now explore prototype two-family models which satisfy the theoretical guidelines outlined above. The main assumption is the presence of an $SO(2)$ family symmetry which

⁵Note that this also holds for models based on rational hierarchy [20].

⁶This result not only satisfies eq. (2.7), but the stronger condition that $\mathcal{M}_D \mathcal{M}_D^T \propto 1$.

acts on the following linear combinations of the current eigenstates⁷ ψ and ψ' (which have the same Standard Model quantum numbers)

$$\psi_{\pm} \equiv \frac{1}{\sqrt{2}}(\psi \pm i\psi'), \quad (3.1)$$

according to

$$\psi_{\pm} \rightarrow e^{\pm i\alpha} \psi_{\pm}. \quad (3.2)$$

We avoid labeling the current eigenstates numerically to avoid confusion with the mass eigenstates $\psi^{(m)}$, which are related to the current eigenstates $\psi^{(c)}$ by

$$\psi^{(c)} \equiv \begin{pmatrix} \psi \\ \psi' \end{pmatrix} = \mathcal{U} \psi^{(m)} \equiv \mathcal{U} \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix}. \quad (3.3)$$

In the current basis, the Dirac mass matrices for fermions of charge q are

$$\mathcal{M}^{(q)} = \mathcal{U}_q^T \mathcal{D}_q \mathcal{V}_q, \quad (3.4)$$

with left- and right-handed mixings \mathcal{U}_q and \mathcal{V}_q , respectively. \mathcal{D}_q is the diagonal matrix of mass eigenvalues. The seesaw matrix \mathcal{M}_{ν} for the light neutrinos is

$$\mathcal{M}_{\nu} = \mathcal{U}_{\nu}^T \mathcal{D}_{\nu} \mathcal{U}_{\nu}, \quad (3.5)$$

and hence in our conventions, the lepton mixing matrix is $\mathcal{U}_{\text{MNSP}} = \mathcal{U}_{-1} \mathcal{U}_{\nu}^{\dagger}$.

***SO*(2) Symmetry limit**

Let us begin with the neutral lepton sector. The *SO*(2) symmetry operates on the lepton doublets L_{\pm} and the right-handed neutrinos \bar{N}_{\pm} , which then enforces degenerate right-handed neutrino masses:

$$2M \bar{N}_+ \bar{N}_- = M (\bar{N} \bar{N} + \bar{N}' \bar{N}'),$$

so that the Majorana mass matrix \mathcal{M}_{Maj} is proportional to the identity matrix. The *SO*(2)-invariant neutral Dirac mass terms are

$$a \nu_+ \bar{N}_- + b \nu_- \bar{N}_+,$$

such that in the current basis the neutral Dirac mass matrix $\mathcal{M}_D^{(0)}$ is

$$\mathcal{M}_D^{(0)} = \begin{pmatrix} z & -iw \\ iw & z \end{pmatrix}, \quad (3.6)$$

where $z = (a + b)/2$, and $w = (a - b)/2$. Note that this matrix is of the form

$$\mathcal{M}_D^{(0)} = z1 - iw(i\sigma_2), \quad (3.7)$$

⁷All fermions are taken to be left-handed. Barred fields denote the right-handed states.

which satisfies the conditions given in Eqs. (2.7)–(2.9). It has the properties that its eigenvalues are nondegenerate

$$m_+^D = |z + w| = |a|, \quad m_-^D = |z - w| = |b|, \quad (3.8)$$

and the left-handed fields are maximally mixed:

$$\mathcal{U}_0 = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & -i \\ 1 & i \end{pmatrix} = \mathcal{R}(-\pi/4)\mathcal{P}, \quad (3.9)$$

in which $\mathcal{R}(\theta)$ denotes a (2×2) rotation by θ and \mathcal{P} is a diagonal phase matrix:

$$\mathcal{P} = \begin{pmatrix} 1 & 0 \\ 0 & i \end{pmatrix}. \quad (3.10)$$

However, the *physical* neutrino masses $m_{\nu_{1,2}}$ are degenerate:

$$\mathcal{M}_\nu = \mathcal{M}_D^{(0)} \mathcal{M}_{\text{Maj}}^{-1} (\mathcal{M}_D^{(0)})^T = \frac{1}{M} \begin{pmatrix} z^2 - w^2 & 0 \\ 0 & z^2 - w^2 \end{pmatrix}, \quad (3.11)$$

and there is no observable mixing. The $SO(2)$ symmetry requires both the right-handed and the left-handed (through the seesaw) neutrino masses to be fully degenerate, even though the Dirac masses can be hierarchical (one can even vanish if $z = \pm w$). For real z and w , a useful parametrization is

$$z = \mu \cosh \eta, \quad w = \mu \sinh \eta,$$

where μ is a mass parameter; alternatively

$$a = \mu e^\eta, \quad b = \mu e^{-\eta},$$

suggestive of Yukawa couplings proportional to warp factors.

Turning to the charged lepton sector, it is simplest to assume that under the family symmetry, the right-handed charged leptons transform as

$$\bar{e}_\pm \rightarrow e^{\pm i\alpha} \bar{e}_\pm. \quad (3.12)$$

$SO(2)$ invariance yields the charged lepton Yukawa couplings

$$f_e e_+ \bar{e}_- + h_e e_- \bar{e}_+,$$

so that in the current basis, the charged lepton Dirac mass matrix is

$$\mathcal{M}_D^{(-1)} = \begin{pmatrix} u & -iv \\ iv & u \end{pmatrix}, \quad (3.13)$$

where $u = (f_e + h_e)/2$ and $v = (f_e - h_e)/2$. The analysis proceeds in a similar way to that of the neutral Dirac masses. The mass eigenvalues are given by

$$m_{e_1} = |u + v| = |f_e|, \quad m_{e_2} = |u - v| = |h_e|. \quad (3.14)$$

If the Yukawa couplings are real we can use the same parametrization, and set

$$f_e = \mu e^{-\eta_e}, \quad h_e = \mu e^{\eta_e}. \quad (3.15)$$

The diagonalization produces the same type of left-handed mixing matrix, *i.e.*

$$\mathcal{U}_{-1} = \mathcal{R}(-\pi/4)\mathcal{P}. \quad (3.16)$$

Even though the charged lepton mass matrix is diagonalized by a maximal 45° rotation, lepton mixing is unobservable because of the neutrino mass degeneracy.

If the neutrinos had only Dirac mass terms, the large angle rotation of the charged lepton sector would be completely undone by the large angle rotation of the neutrino sector in the symmetry limit. Symmetry breaking in the $\Delta I_w = 1/2$ sector would yield a nonvanishing mixing angle, but since this sector is non-degenerate, the mixing would be proportional to the symmetry breaking parameter and therefore small, in contradiction with the neutrino data. The seesaw mechanism is therefore crucial in leading to the possibility of large observable mixing from small symmetry breaking effects.

Finally we note that we could have considered different family charge assignments for the right-handed charged leptons. There is a special case where \bar{e} and \bar{e}' have the same $SO(2)$ charge:

$$\bar{e} \rightarrow e^{-i\alpha} \bar{e}; \quad \bar{e}' \rightarrow e^{-i\alpha} \bar{e}', \quad (3.17)$$

i.e., opposite of L_+ . The allowed charged lepton mass terms are

$$e_+(c\bar{e} + d\bar{e}').$$

The coupling to the other combination e_- is forbidden by the symmetry, or relatively suppressed, as it would be in the Froggatt-Nielsen approach. The charged lepton mass matrix is now of a different form:

$$\mathcal{M}_D^{(-1)} = \frac{1}{\sqrt{2}} \begin{pmatrix} c & d \\ ic & id \end{pmatrix}. \quad (3.18)$$

The charged fermion masses are

$$m_{e_1} = 0, \quad m_{e_2} = \sqrt{c\bar{c} + d\bar{d}}.$$

The left-handed mixings are again given by

$$\mathcal{U}_{-1} = \mathcal{R}(-\pi/4)\mathcal{P}. \quad (3.19)$$

Again, although the charged lepton mass matrix is diagonalized by a maximal 45° rotation, it yields no observable mixing due to the neutrino mass degeneracy.

This charge assignment does not lead to qualitatively different physics from the previous case, except that one of the charged leptons is massless. However, it is not as natural as the previous charge assignment if we assign the same $SO(2)$ charges to complete grand-unified families.

We assume that the family symmetry acts in the same way on the quarks,

$$\mathbf{u}_\pm \rightarrow e^{\pm i\alpha} \mathbf{u}_\pm, \quad \mathbf{d}_\pm \rightarrow e^{\pm i\alpha} \mathbf{d}_\pm, \quad (3.20)$$

leading to the mass matrices:

$$(f_u \mathbf{u}_+ \bar{\mathbf{u}}_- + h_u \mathbf{u}_- \bar{\mathbf{u}}_+) + (f_d \mathbf{d}_+ \bar{\mathbf{d}}_- + h_d \mathbf{d}_- \bar{\mathbf{d}}_+).$$

The analysis proceeds exactly as before: all four quarks are massive, but since the two Dirac matrices are diagonalized in the same way, the CKM mixing matrix is proportional to the identity (which is not a bad starting point).

To summarize this model in the limit of exact symmetry, we see that all charged particles are massive. Quarks and charged leptons with the same quantum numbers have different masses while the two left-handed (and the two right-handed) neutrinos are mass degenerate. None of these particles mix; both families are stable. Next we explore the effects of family symmetry breaking.

Symmetry breaking

Observable lepton mixings require distinct neutrino masses: our $SO(2)$ family symmetry must be broken. The origin of this breaking is not known, but in principle it can occur both in the $\Delta I_w = 1/2$ and the $\Delta I_w = 0$ sectors.

• $\Delta I_w = 0$ **sector**. If the flavor symmetry breaking lies entirely in the $\Delta I_w = 0$ mass matrix, we can expect new effects only in the neutrino sector; all electroweak doublet Dirac matrices remain unaffected. Since the Majorana matrix is symmetric, the symmetry-breaking terms can occur along σ_3 and/or σ_1 .

Let us begin by considering symmetry breaking along σ_3 , such that

$$\mathcal{M}_{\text{Maj}} = M \begin{pmatrix} 1 + \delta & 0 \\ 0 & 1 - \delta \end{pmatrix}, \quad (3.21)$$

where $\delta \ll 1$ is the symmetry-breaking parameter. The degeneracy of the right-handed neutrinos is broken, yielding the seesaw matrix

$$\mathcal{M}_\nu = \hat{\mu} \begin{pmatrix} 1 - \delta \cosh 2\eta & -i\delta \sinh 2\eta \\ -i\delta \sinh 2\eta & 1 + \delta \cosh 2\eta \end{pmatrix}, \quad (3.22)$$

with $\hat{\mu} \equiv \mu^2 / (M(1 - \delta^2))$. The Hermitian combination

$$\mathcal{M}_\nu \mathcal{M}_\nu^\dagger = \hat{\mu}^2 \begin{pmatrix} 1 - 2\delta \cosh 2\eta + \delta^2 \cosh 4\eta & -i\delta^2 \sinh 4\eta \\ i\delta^2 \sinh 4\eta & 1 + 2\delta \cosh 2\eta + \delta^2 \cosh 4\eta \end{pmatrix} \quad (3.23)$$

has distinct eigenvalues:

$$m_{\nu_{1,2}}^2 = \hat{\mu}^2 \left(1 + \delta^2 \cosh 4\eta \mp 2\delta \cosh 2\eta \sqrt{1 + \delta^2 \sinh^2 2\eta} \right). \quad (3.24)$$

The mass-squared difference is therefore of $\mathcal{O}(\delta)$:

$$\Delta m^2 = |m_{\nu_2}^2 - m_{\nu_1}^2| = 4\delta\hat{\mu}^2 \cosh 2\eta \sqrt{1 + \delta^2 \sinh^2 2\eta}. \quad (3.25)$$

For small δ , the left-handed mixing matrix takes a nearly diagonal form

$$\mathcal{U}_\nu = \begin{pmatrix} 1 & -\frac{i\delta}{2} \sinh 2\eta \\ -\frac{i\delta}{2} \sinh 2\eta & 1 \end{pmatrix} + \mathcal{O}(\delta^2), \quad (3.26)$$

providing a small shift from the maximal mixing angle generated in the charged lepton sector:

$$\mathcal{U}_{\text{MNSP}} = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 + \frac{\delta}{2} \sinh 2\eta & -(1 - \frac{\delta}{2} \sinh 2\eta) \\ 1 - \frac{\delta}{2} \sinh 2\eta & 1 + \frac{\delta}{2} \sinh 2\eta \end{pmatrix} \mathcal{P} + \mathcal{O}(\delta^2). \quad (3.27)$$

Family symmetry breaking is seen to generate an observable MNSP angle:

$$\theta = \frac{\pi}{4} - \frac{\delta \sinh 2\eta}{2} + \mathcal{O}(\delta^2), \quad (3.28)$$

hazed away from 45° by a slightly broken continuous family symmetry. As anticipated, degenerate perturbation can create a large angle; in this case the small parameter enters as deviation from $\pi/4$.

Let us now consider symmetry breaking along σ_1 , for which

$$\mathcal{M}_{\text{Maj}} = M \begin{pmatrix} 1 & \delta \\ \delta & 1 \end{pmatrix}, \quad (3.29)$$

which leads to the seesaw mass matrix

$$\mathcal{M}_\nu = \hat{\mu} \begin{pmatrix} 1 + i\delta \sinh 2\eta & -\delta \cosh 2\eta \\ -\delta \cosh 2\eta & 1 - i\delta \sinh 2\eta \end{pmatrix}. \quad (3.30)$$

The diagonal elements of the Hermitian combination $\mathcal{M}_\nu \mathcal{M}_\nu^\dagger$ are degenerate:

$$\mathcal{M}_\nu \mathcal{M}_\nu^\dagger = \hat{\mu}^2 \begin{pmatrix} 1 + \delta^2 \cosh 4\eta & -2\delta \cosh 2\eta (1 + i\delta \sinh 2\eta) \\ -2\delta \cosh 2\eta (1 - i\delta \sinh 2\eta) & 1 + \delta^2 \cosh 4\eta \end{pmatrix}, \quad (3.31)$$

such that the left-handed mixings are given by

$$\mathcal{U}_\nu = \mathcal{P}_\nu \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & e^{-i\rho} \\ -e^{i\rho} & 1 \end{pmatrix}, \quad (3.32)$$

which describes a maximal mixing angle. In the above,

$$\tan \rho = \delta \sinh 2\eta, \quad (3.33)$$

and the nontrivial phase matrix is

$$\mathcal{P}_\nu = \begin{pmatrix} e^{i\rho/2} & 0 \\ 0 & e^{-i\rho/2} \end{pmatrix}. \quad (3.34)$$

The eigenvalues are the same as for diagonal breaking, and once again

$$\Delta m^2 = |m_{\nu_2}^2 - m_{\nu_1}^2| = 4\delta\hat{\mu}^2 \cosh 2\eta \sqrt{1 + \delta^2 \sinh^2 2\eta} . \quad (3.35)$$

Although \mathcal{U}_{-1} and \mathcal{U}_ν contain 45° rotations, the MNSP matrix takes the form

$$\mathcal{U}_{\text{MNSP}} = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{1 + \sin \rho} & -e^{-i\rho} \sqrt{1 - \sin \rho} \\ e^{i\rho} \sqrt{1 - \sin \rho} & \sqrt{1 + \sin \rho} \end{pmatrix} \mathcal{P}\mathcal{P}_\nu^* + \mathcal{O}(\delta^2) , \quad (3.36)$$

which, for small δ , describes a rotation angle θ , slightly hazed away from 45° :

$$\tan \theta = \sqrt{\frac{1 - \sin \rho}{1 + \sin \rho}} \approx 1 - \delta \sinh 2\eta , \quad (3.37)$$

such that as in the case of diagonal breaking,

$$\theta = \frac{\pi}{4} - \frac{\delta \sinh 2\eta}{2} + \mathcal{O}(\delta^2) . \quad (3.38)$$

It can be shown that symmetry breaking along σ_1 and σ_3 are identical up to phases in the MNSP matrix. For off-diagonal breaking, nontrivial phases prevent a cancellation of the large mixings of the charged and neutral leptons, resulting in a nearly maximal observable mixing angle. These phases include the observable CP-violating Majorana phase present in the two family case.

• $\Delta I_w = 1/2$ **sector**. As previously stated, quark mixings can arise only if the family symmetry is broken in the electroweak doublet sector. In principle such breaking can be generated through radiative effects stemming from the breaking in the $\Delta I_w = 0$ sector, but for the present purposes, let us assume that symmetry breaking effects arise solely in the $\Delta I_w = 1/2$ sector of the theory. Such effects can occur along with the $SO(2)$ invariant Yukawa couplings as follows:

$$a\nu_+ \bar{N}_- + b\nu_- \bar{N}_+ + \mu\delta\nu_+ \bar{N}_+ + \mu\delta\nu_- \bar{N}_- , \quad (3.39)$$

where we have assumed that the breaking occurs in the $|\Delta F = 2|$ terms only, with strength δ . In the current basis, the Dirac mass matrix now becomes

$$\mathcal{M}_D = \mu \begin{pmatrix} \cosh \eta + \delta & -i \sinh \eta \\ i \sinh \eta & \cosh \eta - \delta \end{pmatrix} . \quad (3.40)$$

The seesaw matrix then takes the form

$$\mathcal{M}_\nu = \mathcal{M}_D \mathcal{M}_D^T = \frac{\mu^2}{M} \begin{pmatrix} 1 + 2\delta \cosh \eta + \delta^2 & 2i\delta \sinh \eta \\ 2i\delta \sinh \eta & 1 - 2\delta \cosh \eta + \delta^2 \end{pmatrix} . \quad (3.41)$$

To leading order in δ , the masses are

$$m_{1,2}^2 = \frac{\mu^2}{M} (1 \mp 4\delta \cosh \eta + \mathcal{O}(\delta^2)) , \quad (3.42)$$

and the mixing matrix is

$$\mathcal{U}_\nu = \begin{pmatrix} 1 & i\delta \sinh \eta \\ i\delta \sinh \eta & 1 \end{pmatrix} + \mathcal{O}(\delta^2) . \quad (3.43)$$

Hence, the mass-squared difference and the MNSP mixing angle are

$$\Delta m^2 = 8\delta \cosh \eta + \mathcal{O}(\delta^2), \quad \theta = \frac{\pi}{4} + \delta \sinh \eta + \mathcal{O}(\delta^2). \quad (3.44)$$

For the neutrinos, the results are essentially the same as for electroweak-singlet breaking, with the seesaw as the crucial ingredient for generating large mixing.

This can best be seen by considering the effect of this breaking in the quark sector. We take the quark mass matrices to be

$$m_u(e^{-i\eta_u} \mathbf{u}_+ \bar{\mathbf{u}}_- + e^{i\eta_u} \mathbf{u}_- \bar{\mathbf{u}}_+) + m_u \delta_u (\mathbf{u}_+ \bar{\mathbf{u}}_+ + \mathbf{u}_- \bar{\mathbf{u}}_-), \quad (3.45)$$

and

$$m_d(e^{-i\eta_d} \mathbf{d}_+ \bar{\mathbf{d}}_- + e^{i\eta_d} \mathbf{d}_- \bar{\mathbf{d}}_+) + m_d \delta_d (\mathbf{d}_+ \bar{\mathbf{d}}_+ + \mathbf{d}_- \bar{\mathbf{d}}_-). \quad (3.46)$$

As a result of the perturbation ($\delta \ll 1$), the masses are shifted to

$$m_{u_1} = m_u \left(e^{-\eta_u} - \frac{\delta_u^2}{2 \sinh \eta_u} \right); \quad m_{u_2} = m_u \left(e^{\eta_u} + \frac{\delta_u^2}{2 \sinh \eta_u} \right), \quad (3.47)$$

and similarly for the down quarks with $u \rightarrow d$. The left-handed mixing matrices are simply

$$\mathcal{U}_{2/3} = \begin{pmatrix} \cos \theta_u & \sin \theta_u \\ -\sin \theta_u & \cos \theta_u \end{pmatrix} \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & i \\ 1 & -i \end{pmatrix}, \quad (3.48)$$

with

$$\sin \theta_u \approx \frac{\delta_u}{2 \sinh \eta_u}, \quad (3.49)$$

while

$$\mathcal{U}_{-1/3} = \begin{pmatrix} \cos \theta_d & \sin \theta_d \\ -\sin \theta_d & \cos \theta_d \end{pmatrix} \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & i \\ 1 & -i \end{pmatrix}, \quad (3.50)$$

with

$$\sin \theta_d \approx \frac{\delta_d}{2 \sinh \eta_d}. \quad (3.51)$$

It follows that the CKM rotation angle is

$$\theta = \theta_u - \theta_d \approx \frac{\delta_u}{2 \sinh \eta_u} - \frac{\delta_d}{2 \sinh \eta_d}. \quad (3.52)$$

In the limit of strong hierarchies, $\eta_{u,d} \gg 1$, we can write the mixing angle as

$$\theta \approx \delta_u \sqrt{\frac{m_{u_1}}{m_{u_2}}} - \delta_d \sqrt{\frac{m_{d_1}}{m_{d_2}}}. \quad (3.53)$$

Unfortunately, we cannot estimate its value since $\delta_{u,d}$ are arbitrary.

These instructive two-family models demonstrate how large angles can arise from a *slightly broken continuous family symmetry*. With exact symmetry, the $\Delta I_w = 1/2$ masses are hierarchical and the light neutrinos are mass-degenerate after the seesaw, even with degenerate $\Delta I_w = 0$ masses. The symmetry breaking schemes yield large lepton mixings, with small deviations from maximal mixing. The approach lends itself to a further embedding in a grand unified picture, as the structure of the $\Delta I_w = 1/2$ mass matrices yields a good theoretical starting point for describing the hierarchical quark and charged lepton masses.

We note here that the idea of obtaining large angles using $U(1)$ family symmetries has been previously explored in the literature; an explicit example is the two-family scenario of [21]. However, our scenario has several important differences from this model and most other Abelian flavor models. First, there is a clear difference in the philosophy of our approach, which is to motivate symmetries by requiring degeneracies, rather to use the usual flavor model-building methodology, which is to demand a family symmetry and explore its resulting implications. This difference, though subtle, is one which does motivate the particular two-family $SO(2)$ scenarios presented here, which display a certain simplicity and elegance (e.g. with simple charge assignments). The $SO(2)$ models also have an intriguing potential embedding in the braneworld context, given the hierarchical Yukawa couplings required at Froggatt-Nielsen tree level.⁸ However, the most important difference is the requirement of *complex* Dirac mass matrices, which singles out a *nonstandard flavor basis with nontrivial phases* (the $\psi \pm i\psi'$ combination of fields as opposed to the standard ψ, ψ' basis). In most $U(1)$ flavor models, there is no natural explanation for these phases; they are just part of the usual $O(1)$ coefficients which Abelian symmetries do not fix. For this reason, these phases, which play a crucial role in our approach) are generically not found in [21] or in other Abelian flavor models in the literature.

We next discuss possible generalizations of this mechanism to three families.

4. Extensions to three families

For three families, the naive generalization of the $O(2)$ symmetry which naturally emerges in the two-family case to $O(3)$ does not allow for hierarchical Dirac mass matrices and degenerate neutrinos along the lines of our approach, as discussed in section 2. The question of interest, therefore, is how to generalize the elegant mechanism of the two-family model based on $SO(2)$ to the case of three-family mixing. The main theoretical challenge is to reproduce (or more optimistically, explain) the intriguing pattern of mixing angles seen in the data.

Although a full $O(3)$ symmetry does not satisfy the requirements of our theoretical approach, there are other ways to generalize the approach to three family models. One way is to keep the same $SO(2)$ symmetry and simply assign family charges to the third family fields in a way which incorporates the two-family mixing results. Another is to

⁸Note that higher-order FN effects or symmetry breaking dynamics are often required to achieve hierarchies in flavor models, but are not needed here (at least not at leading order). A complete exploration of these effects is beyond the scope of this paper.

seek alternate symmetries which may or may not include the family charge of the previous section.

SO(2) Models

We begin with the $SO(2)$ models. In this case, we simply assign $SO(2)$ charges to the fields of a third family, L'' , \bar{e}'' , \bar{N}'' , \mathbf{Q}'' , $\bar{\mathbf{u}}''$, and $\bar{\mathbf{d}}''$. One possibility is that the third family fermions are uncharged under the $SO(2)$. If, on the other hand, the third family fields have nonvanishing $SO(2)$ charges, each third family field ψ'' either has the same charge as ψ_+ or ψ_- (and mix with these states in the symmetry limit), *i.e.*:

$$\psi'' \rightarrow e^{\pm i\alpha} \psi''.$$

In what follows, we will provide a brief summary of these two situations, relegating details and explicit models to a more comprehensive analysis given in [22].

• **Vanishing third family $SO(2)$ charges.**

When the third family is blind to the $SO(2)$ symmetry, we will see that the path to degeneracy is arguably the most straightforward. For the neutral leptons, it is straightforward to see that degenerate right-handed neutrinos arise in the limit of exact family symmetry when \bar{N}'' and L'' are uncharged under the $SO(2)$:

$$2M\bar{N}_+\bar{N}_- + M_0\bar{N}''\bar{N}'' \tag{4.1}$$

$M_0 = M$ corresponds to $SO(3)$ invariance in the electroweak singlet sector, which could arise from a custodial symmetry which is present in this sector but not in the electroweak doublet sector. The neutral Dirac couplings are

$$\mu e^\eta \nu_+ \bar{N}_- + \mu e^{-\eta} \nu_- \bar{N}_+ + \mu_0 \nu'' \bar{N}'' \tag{4.2}$$

The Dirac mass eigenvalues are nonvanishing and hierarchical. In the symmetry limit, all three neutrinos seesaw:

$$m_{1,2} = \frac{\mu^2}{M}, \quad m_3 = \frac{\mu_0^2}{M_0}.$$

Two of the masses are naturally degenerate, but total degeneracy can be achieved depending on the parameters of the theory. The mixing angles and detailed mass splittings then depend on the form of the symmetry breaking, which can occur in the $\Delta I_w = 0$ and/or $\Delta I_w = 1/2$ sectors. For example, in the right-handed neutrino sector one can have symmetry breaking terms which couple \bar{N}'' to \bar{N}_\pm (*i.e.* $\Delta F = \pm 1$ effects, in which F denotes the family charge), or corrections to the couplings of the \bar{N}_\pm sector ($\Delta F = 2$ effects). The main theoretical challenge in both cases is to achieve the two large mixing angles while keeping θ_{13} small, which constrains the allowed symmetry breaking effects.

• **Nonvanishing third family $SO(2)$ charges.**

In this case, the third family fields are charged under the $SO(2)$ and mix with the other families in the symmetry limit. For reasons which will become clear shortly, we parametrize this mixing as

$$\begin{aligned} \psi_\pm(\zeta, \phi) &= \cos \zeta \psi_\pm + \sin \zeta e^{i\phi} \psi'' \\ \psi_\pm^\perp(\zeta, \phi) &= -\sin \zeta e^{-i\phi} \psi_\pm + \cos \zeta \psi'', \end{aligned} \tag{4.3}$$

in which ζ and ϕ are arbitrary angles. Since certain Standard Model fields enter more than one type of mass term (e.g. lepton doublets enter both the charged and neutral Dirac mass terms, and right-handed neutrinos enter both the neutral Dirac and Majorana masses), when the third family leptons have $SO(2)$ charges, the phenomenological discussion can be cast into the question of which linear combinations of fields enter the different types of mass terms.

When the *same* linear combinations enter two types of Yukawa couplings, we say that they are “locked.” A familiar example of near locking is that of the quark doublets, which enter the up-type and down-type Yukawa couplings in nearly the same way up to Cabibbo-sized effects; *i.e.*, the CKM matrix measures the deviation from perfect locking. For the lepton sector, if the seesaw mechanism is operational the observed large angles of the MNSP matrix do not preclude the possibility of locking in either the L or \bar{N} sectors.

Locking can be achieved by invoking a second family phase symmetry $U(1)_\beta$, under which e.g. for the \bar{N} sector (assuming for concreteness that \bar{N}'' and \bar{N}_- have the same charges under the original $SO(2)_\alpha$):

$$(\bar{N}_- - i\bar{N}'') \rightarrow e^{-i\beta}(\bar{N}_- - i\bar{N}''), \quad (\bar{N}_- + i\bar{N}'') \rightarrow e^{i\beta}(\bar{N}_- + i\bar{N}''),$$

which requires the same linear combination of \bar{N} 's to enter the $\Delta I_w = 0$ and the $\Delta I_w = 1/2$ couplings. Similar considerations apply to the lepton doublets.

With this in mind, note that there are two coupling schemes of interest, depending on the relative signs of the charges of L'' and \bar{N}'' :

- *Type A couplings.* In this coupling scheme, L'' and \bar{N}'' have the same nonvanishing $SO(2)$ charge; we assume for concreteness that they mix with L_- and \bar{N}_- , respectively. The $SO(2)$ invariant $\Delta I_w = 0$ mass terms are then

$$M\bar{N}_+\bar{N}_-(\zeta, \phi). \tag{4.4}$$

As only one combination of \bar{N}_- and \bar{N}'' enters in the Majorana mass terms, the orthogonal combination $\bar{N}_-^\perp(\zeta, \phi)$ does not participate in the seesaw. Similarly, it is straightforward to see that without loss of generality, the $SO(2)$ invariant neutral $\Delta I_w = 1/2$ mass terms are of the form

$$\mu e^\eta \nu_+ \bar{N}_-(\zeta', \phi') + \mu e^{-\eta} \nu_-(\chi, \xi) \bar{N}_+. \tag{4.5}$$

As only one combination of \bar{N} 's and L 's enters the Dirac mass terms, $\nu^\perp(\chi, \xi)$ and $\bar{N}^\perp(\zeta, \phi)$ remain massless in the $SO(2)$ symmetry limit. The physics also depends on whether there is locking. If locking takes place in the \bar{N} sector, $\zeta = \zeta'$ and $\phi = \phi'$. Alternatively, there could be locking in the L sector, such that the combination $e_-(\chi', \xi')$ which enters the charged lepton couplings (which depends on the charges of the $\bar{\nu}$'s; there is always an orthogonal combination $e_-^\perp(\chi', \xi')$ which does not enter) satisfies the conditions $\chi = \chi'$ and $\xi = \xi'$.

With or without locking, the pattern of masses and mixings in the limit of exact $SO(2)$ symmetry is unrealistic, but this can be remedied by symmetry breaking. If symmetry breaking occurs only in the $\Delta I_w = 0$ sector, the absence of ν^\perp in the Yukawa couplings implies one massless left-handed neutrino: small perturbations (minimally to give \bar{N}_-^\perp

a large Majorana mass) will produce the so-called inverted hierarchy. In this case, the neutrino mass splitting is related to the haze away from maximal mixing, and applies to solar oscillations. The haze turns out to be too small to explain the solar angle, although it naturally yields $\theta_{13} = 0$ in the symmetry limit. The atmospheric angle is not determined, but instead can be directly traced to the absence of locking in the L sector (*i.e.*, $\theta_{\oplus} \sim \chi' - \chi$). While these minimal models are not consistent with the data, more promising models can be constructed by including $\Delta I_w = 1/2$ symmetry breaking couplings that bring ν^\perp into the neutrino seesaw.

- *Type B couplings.* In this class of models, L'' and \overline{N}'' have nonvanishing but opposite $SO(2)$ charges; e.g., let us assume for concreteness that \overline{N}'' mixes with \overline{N}_- and L'' mixes with L_+ . In this case, the $\Delta I_w = 0$ mass terms are given by eq. (4.4) as before. However, the $\Delta I_w = 1/2$ couplings are different:

$$\mu e^\eta \nu_+(\chi, \xi) \overline{N}_-(\zeta', \phi') + \mu e^{-\eta} \nu_-(\chi, \xi) \overline{N}_+(\zeta', \phi') + \mu' \nu_+^\perp(\chi, \xi) \overline{N}_-^\perp(\zeta', \phi'). \quad (4.6)$$

In the above, μ' is assumed to be of the order of the electroweak scale (*i.e.*, μ' is of $\mathcal{O}(\mu)$, as motivated by the logarithmic spacing of the quark and charged lepton masses), but otherwise is a free parameter. Since both ν^\perp and \overline{N}^\perp have Yukawa couplings, the neutral Dirac matrix has no zero eigenvalues. Furthermore, there are no massless neutrinos in the $SO(2)$ symmetry limit. Two of the neutrinos have seesaw suppressed masses, while the other two neutrinos generically have electroweak scale masses, whether or not there is \overline{N} locking.

The neutrino masses and mixings are again not realistic in the symmetry limit for Type B couplings. With $\Delta I_w = 0$ symmetry breaking terms (again minimally to give \overline{N}^\perp a large Majorana mass), the phenomenological implications of the resulting models depend strongly on whether \overline{N} locking occurs or not. With \overline{N} locking, θ_{13} is naturally zero (which generically occurred for Type A models) independently of the form of the symmetry breaking terms in the $\Delta I_w = 0$ sector. However, the Type B models also do not require an inverted hierarchy scheme, and hence lead to a greater flexibility for constructing viable models. For example, models can be constructed with minimal tuning in which the two-family mechanism described in section 3 can be successfully applied to atmospheric oscillations; the solar angle is then directly related to the absence of locking in the L sector. Without \overline{N} locking, there is again a greater flexibility to construct models due to the various possibilities for the mass hierarchy. However, in this case generically $\theta_{13} \neq 0$, and hence the theoretical challenge is to obtain the two large mixings while keeping θ_{13} small.

For both Type A and B couplings, only two right-handed neutrinos are degenerate; *i.e.*, $\mathcal{M}_{\text{Maj}} \neq 1$ in the symmetry limit. Hence, when the third family fields are charged under the original $SO(2)$ symmetry, the resulting models step outside of our theoretical framework. For Type A models, light neutrino mass degeneracy will not occur unless there is a conspiracy between specific symmetry breaking terms in both the $\Delta I_w = 0$ mass terms (to give the third right-handed neutrino a mass) and $\Delta I_w = 1/2$ couplings (to involve the third family in the seesaw). For Type B couplings, degeneracy of the light neutrinos requires symmetry breaking in the $\Delta I_w = 0$ sector and a specific tuning of the parameters.

In summary, the three-family $SO(2)$ schemes provide an interesting framework for model building, but have the negative generic feature that neutrino mass degeneracy is lost (for both light and heavy neutrinos). The lack of degeneracy is also a hallmark of more general attempts to obtain viable three-family mixing schemes with Abelian family symmetries. Despite this shortcoming, our approach may provide new insights or arise in particular settings. Further exploration of these scenarios will be presented in [22].

Outlook: the path to neutrino mass degeneracy

We have shown in this paper that for two families, it is possible to obtain neutrino mass degeneracy (via the seesaw) through a continuous family symmetry, while allowing for hierarchical charged fermion masses. However, as discussed in section 2, generalization to three families has proven more challenging. Let us recall the constraints imposed on the Dirac and Majorana matrices by requiring for simplicity that the seesaw combination

$$M_D \frac{1}{\mathcal{M}_{\text{Maj}}} M_D^T \tag{4.7}$$

is proportional to the identity matrix (a stronger condition than was previously imposed in section 2). If we write the Dirac matrix in terms of three vectors $\mathbf{v}_{[i]}$, we can think of the inverse Majorana matrix as a metric in this three dimensional space. In its “flat” limit, the vectors are orthogonal

$$\mathbf{v}_{[i]}^T \mathbf{v}_{[j]} \sim \delta_{ij} . \tag{4.8}$$

If the vector entries were all real, the Dirac matrix would simply be a rotation matrix, in which case its three eigenvalues would be equal, and would not produce any hierarchy. In the (2×2) case, we could produce a hierarchy by introducing a family $SO(2)$, but then the vectors had complex entries.

In the (3×3) case, a matrix that fulfills our requirement is

$$\mathcal{M}_D = \begin{pmatrix} \cosh \eta & i \sinh \eta \cos \theta & i \sinh \eta \sin \theta \\ -i \sinh \eta & \cosh \eta \cos \theta & \cosh \eta \sin \theta \\ 0 & -\sin \theta & \cos \theta \end{pmatrix} . \tag{4.9}$$

Its eigenvalues $(1, e^\eta, e^{-\eta})$ show equal logarithmic spacing of the masses, clearly a desirable feature for charged fermion Dirac matrices. Unfortunately, we have not found a symmetry principle from which this naturally emerges.

A second approach is to invoke non-Abelian family symmetries. For three families, the splitting of the neutrino masses from full degeneracy can be viewed as a small λ_8 breaking and smaller λ_3 breaking, suggesting an $SU(3)$ family symmetry, or a small breaking along m^2 and smaller breaking along m , where m is the flavor magnetic quantum number of an $SO(3)$ family symmetry.

We have already discussed the shortcomings of a full family $SO(3)$ here. A family $SU(3)$ can naturally describe the splittings, but the path to degeneracy is again quite different than our approach. With one right-handed neutrino per family, the Majorana mass matrix transforms as a family sextet and can be highly hierarchical (see e.g. [23, 24]).

However, we can augment the number of \overline{N} 's: for example, E_6 requires two right-handed neutrinos per family.

In such models, an amusing possibility is to invoke an additional phase symmetry which allows the six neutrinos to be arranged into one triplet and one antitriplet. Invariance under this phase symmetry forces their mass matrix to be either a family singlet or a family octet. In the limit of exact symmetry, the three left-handed neutrinos stay massless, achieving degeneracy. After symmetry breaking, they acquire masses through the seesaw. The light neutrino mass splittings can then be described by perturbations along λ_8 and λ_3 . Although this deviates from the specific theoretical approach of this paper, it is an intriguing path to degeneracy which may provide a fruitful ground for model building. We hope to return to these issues elsewhere.

In summary, we have outlined a flavor model building framework in which continuous family symmetries emerge from the theoretical requirements of hierarchical charged fermion masses and degenerate neutrino masses for both heavy and light neutrinos. Neutrino mass degeneracy is the starting point because large mixings can arise from small perturbations about degenerate structures, and quark-lepton unification suggests that such small parameters should be expected in the lepton sector as well as in the quark sector.

The requirements of degeneracy with hierarchy naturally stem from symmetry for two families, but for three families this does not naturally emerge and finding a symmetry-based starting point is more difficult. However, one should keep in mind that it is a formidable challenge to explain the three-family lepton mixing pattern together with the quarks in a truly satisfactory way, which is why the flavor puzzle has taken such an intriguing turn in light of the recent lepton data. Although there are challenges in extending our approach to three families which mirror those found in general $U(1)$ models, the approach is incremental, and it does show what we must do to achieve naturalness in a three family context. The work therefore provides a particular setting in which to investigate these issues which may yield insight or naturally occur in specific theoretical contexts, and as such is one of many possible avenues worthy of further exploration in addressing the flavor puzzle of the Standard Model.

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